

# What Volume Should the Probe Scan?

The size of the volume scanned by the probe is proportional to the spacing  $\rho$  between the emitting fiber and the detector. The spacing must be adapted to the physical and optical properties of the scanned medium. For the study of sea ice, the scanned volume could be in the order of the cm<sup>3</sup> for two reasons:

1) Brines, salts, bubbles, algae and ice crystals have size distributions ranging from less than a micron to centimeters. Therefore, the scanned volume must be in

2) As the source-detector distance p decreases, more terms N need to be included in the description of the radiance and the phase function to correctly simulate how light is backscattered. At N=2, the modified Henyey-Greenstein relation can be used to describe angular dependence of light after a scattering event

 $p_{mHG}(\boldsymbol{\Theta}) = \propto \cdot \frac{1}{4\pi} \cdot \frac{1 - g^2}{\left(1 + g^2 - 2g\cos\boldsymbol{\Theta}\right)^{3/2}} + (1 - \propto) \cdot \frac{3}{4\pi} \cdot (\cos\boldsymbol{\Theta})^2$ 

Where  $p_{mHG}$  is the probability that a photon will be deviated of an angle  $\theta$ .  $\propto$  and g are 2 terms ranging from 0 to 1 that allow to describe the angular dependence of the photon. For this phase function to be valid, the source-detector spacing  $\rho$  must be bigger than 0.5/b(1-g) according to Bevilacqua (1998) (see table 1). Therefore, the scanned volume should be in the order of the  $cm^3$  to accurately infer the IOPs of the 2 first layers.

> **Table 1.** Estimated Inherent optical properties and minimum spacing between source and detector to use the modified Henyey-Greenstein approximation for the different optical layers of sea ice (Bevilacqua 1998, Ehn 2008, Light, 2008, Light 2015).

Optical layer	a(λ) (m <sup>-1</sup> )	b (m <sup>-1</sup> )	g (H-G) (-)	ρ <sub>min</sub> to use mHG approximation (N=2) (cm)
Surface Scattering Layer	0.1-1	100-1000	0.85	0.33
Drained Layer	0.01-1	10-100	0.85	3.33
Interior Layer	0.01-1	1-10	0.94	83

## **Preliminary Monte Carlo Simulations**

Using the 3D Monte Carlo method to simulate light propagation with a reflectance geometry (see figure 2), 2 preliminary tests were achieved

The scanned depth (see figure 3) allows us to know what volume size we are scanning depending on the distance  $\rho$  from the source and depending on the IOPs of the ice. The Reflectance vs distance from source relation (see figure 4) shows that, for reflectance geometry, using a 2 terms phase function (like modified Henyey-Greenstein) to describe radiance is significant when the source-detector distance is small ( $\rho \sim < 1$  cm).

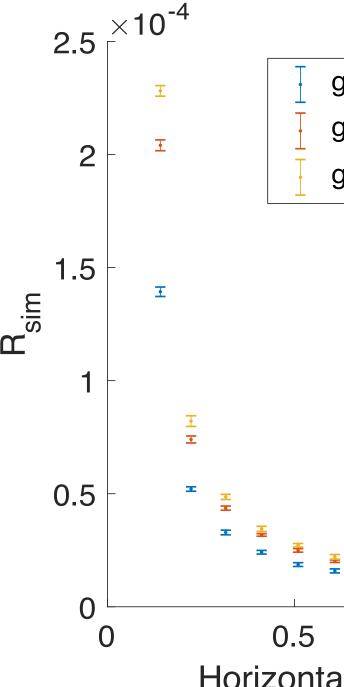
## Next Steps

-Test inversion algorithms: Determine the best algorithm to infer IOPs from field measurements. -Simulate a look-up table: Run simulations varying the phase function asymmetry parameter, absorption and scattering

-Design and build the probe and validate the instrument with phantoms: Build the instrument and validate its functioning

**References:** Bevilacqua, Frédéric. "Local optical characterization of biological tissues in vitro and in vivo." Unpublished PhD diss1781 (1998). Ehn, J., et al. (2008). "Inference of optical properties from radiation profiles within melting landfast sea ice." Journal of Geophysical Research: Oceans 113(C9). Light, B., et al. (2008). "Transmission and absorption of solar radiation by Arctic sea ice during the melt season." Journal of Geophysical Research: Oceans 113(C3). Light, B., et al. (2015). "Optical properties of melting first-year Arctic sea ice." Journal of Geophysical Research: Oceans 120(11): 7657-7675.

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Horizontal distance from source (cm)

**Figure 4.** Simulated reflectance vs distance  $\rho$  from the source varying the second moment (N=2) of the phase function (a=0.1 m<sup>-1</sup> and b=100 m<sup>-1</sup>).



(1)

g<sub>1</sub>=0.85 g<sub>2</sub>=0.7225 (Henyey-Greenstein)  $g_1 = 0.85 g_2 = 0.85$ g<sub>1</sub>=0.85 g<sub>2</sub>=0.9